

STEREO MOC Status Report
Time Period: 2016:095 - 2016:101

STEREO Ahead (STA) Status:

1. The following Ground System anomalies/events occurred during this reporting period:

- On day 095, during the DSS-63 support, turbo decoder lock was lost intermittently beginning at 1813z through 1820z. This anomaly resulted in the loss of 168 frames of real-time and SSR data. See DR# N110270 for more information.
- On day 097, during the DSS-55 support, turbo decoder lock was lost briefly at 1258z. This anomaly resulted in the loss of one frame of SSR data.
- On day 098, during the DSS-25 support, turbo decoder lock was lost briefly at 2331z. This anomaly resulted in the loss of one frame of SSR data.
- On day 099, during the DSS-26 support, turbo decoder lock was lost intermittently beginning at 2322z through 100-0200z. This anomaly resulted in the loss of 31,737 frames of real-time and SSR data. See DR# N110271 for more information.
- On day 100, during the DSS-15 support, turbo decoder lock was lost briefly at 1459z. This anomaly resulted in the loss of one frame of SSR data.
- On day 100, during the DSS-26 support, turbo decoder lock was lost intermittently beginning at 2315z through 101-0003z. This anomaly resulted in the loss of 4072 frames of real-time and SSR data. See DR# N110272 for more information.
- On day 101, the 5th PN ranging demonstration using DSS-15 was cancelled to assist the Kepler mission with recovery from a spacecraft emergency.
- On day 101, during the DSS-25 support, turbo decoder lock was lost briefly at 2030z and again at 2243z. This anomaly resulted in the loss of 2 frames of real-time and SSR data.

2. The following spacecraft/instrument events occurred during this week. The Ahead observatory operated nominally during this week on the center of the HGA main lobe. The HGA feed assembly was at 108 degrees C and decreasing with the HGA angle at 9.6 degrees and increasing, with respect to the spacecraft-Sun line.
 - On day 096, the 33rd SECCHI stepped calibration was executed at 1045z for midpoint in the orbit. This was the 2nd SECCHI stepped calibration roll to be conducted without gyro use. Note that the G&C control bandwidths that were loaded on day 055 for improving roll error were in effect for each data collection step.
 - On day 097, the G&C rate-merge algorithm was enabled in RAM at 1131z. This algorithm was developed post-launch and enables rate measurements from all sensors to be combined to form a "better" rate estimate versus picking only one G&C sensor as the source of the angular rate measurement. While the X-axis jitter has improved, performance throughout a momentum cycle will be monitored. This update is attempting to minimize the increase in roll error that has affected SECCHI HI images intermittently since exiting solar conjunction.
 - The average daily science data return for Ahead was 5.6 Gbits during this week.

STEREO Behind (STB) Status:

1. The following Ground System anomalies/events occurred during this reporting period:
 - None as recovery tracks are every other week.
2. Detailed status of the recovery activities to restore operations from the Behind loss of communication anomaly, which occurred on October 1, 2014, are listed below.
 - The Behind observatory entered superior solar conjunction at the 2.0 degree SPE angle on January 22, 2015. Recovery efforts resumed post solar conjunction on May 4th through June 27, 2015, as the spacecraft had cleared solar interference for LGA communications. The Failure Review Board recommendations were implemented consisting of

battery state of charge recovery and powering on the downlink carrier. The Green Bank Radio Telescope and the Arecibo Observatory also observed the carrier recovery tracks. No downlink signal has been detected. Due to Behind's retrograde motion causing it to re-enter the region of solar interference, recovery operations were suspended from June 28th through November 29, 2015. Weekly recovery operations resumed on November 30, 2015. The Green Bank Radio Telescope and the Allen Telescope Array will also observe the carrier recovery tracks depending on availability. While the Arecibo Observatory is willing also assist, the Behind observatory is not in view until April 2016. Recovery operations were reduced to every other week beginning on March 21st to minimize the impact on DSN resources.

- The Failure Review Board's recommended faster frequency segmented acquisition sequence was tested with the Ahead observatory on September 29, 2015. All 18 one kHz frequency steps were tested twice. While stepping down through the 1 kHz segments, on segment #9 going down in frequency, the transponder locked to the BLF and accepted 9 no-op commands as expected. An interesting finding, but not unexpected, was that the transponder continued to follow the moving carrier and accept all commands sent for the remaining 27 segments.
- As commands must be received to recover the Behind observatory, the first use of the newly developed DSN uplink arraying capability for Behind recovery operations occurred on March 17th. The uplink array capability will provide four times the uplink received power as a 70m station. The uplink array consists of using an 80 kW and two 20 kW transmitters from three 34m Goldstone stations, DSS-24, 25, and 26, with each uplink being precisely phase shifted to create a constructive interference. From testing with STEREO Ahead, an approximate 12 dBm increase in received uplink power, as compared to a single 34m, has been successfully demonstrated each time. As the Behind observatory may be rotating, a frequency segmented acquisition sequence will be used with the uplink array. For each use of the uplink array, as transponder feedback is required, the Ahead observatory is first used to calibrate the phased uplink array then the three stations are switched to point to the Behind observatory sending commands for recovery. The uplink array will be used twice monthly within a three day period, once for battery

recovery and again for carrier recovery. The next use of the uplink array is scheduled for April 16th and 17th.

- As time goes by, the ephemeris error increases degrading DSN antenna pointing. However, with time the spacecraft range also decreases improving RF communications and the ability for other assets to acquire data on Behind. Analysis indicates that the total RF gain change is significant and the probability of command success increases with time.

Significant findings to date:

1. Analysis of the three DSN extracted telemetry frames from the carrier signal just before the planned observatory reset/anomaly occurred on October 1, 2014 showed nominal performance of the spacecraft, i.e., no anomalies, IMU off, and the star tracker providing an attitude solution.
2. Post reset, from the very limited telemetry, three packets, extracted from the carrier signal by the DSN, the X-axis gyro on IMU-A had failed. Unfortunately, this telemetry contained only G&C anomaly data and no spacecraft summary data, i.e., the state of the RF, G&C, fault protection and other subsystems is not known at the time of the anomaly. With a failed IMU and the star tracker being off-line for an undetermined duration, the sun sensors will keep the observatory pointed at the Sun, though the G&C will not have any roll knowledge, and cannot roll the observatory as part of the safing configuration to re-establish communications on the LGAs. From analysis of this telemetry and initial G&C simulations, it is highly suspected that the observatory is rotating about the principal axis of inertia due to an autonomous momentum dump initiated by highly biased gyro data flagged good by the IMU, but this has not yet been confirmed.
3. At least two anomalies occurred post reset, the star tracker not promoting to AAD mode and the X-axis gyro failure. Unfortunately, due to the number of possible combinations, the STEREO fault protection system is not designed for simultaneous failures.

Once communications are restored and the anomaly resolved, the Behind observatory will be returned to nominal science data collection as soon as it is safely possible.